

# **Solar energy**

Purdue ME 597, Distributed Energy Resources

Kevin J. Kircher

# Outline

The solar resource

The sun's position in the sky

Incident irradiance on surfaces

Solar photovoltaics

The sun is (?) a mass of incandescent gas



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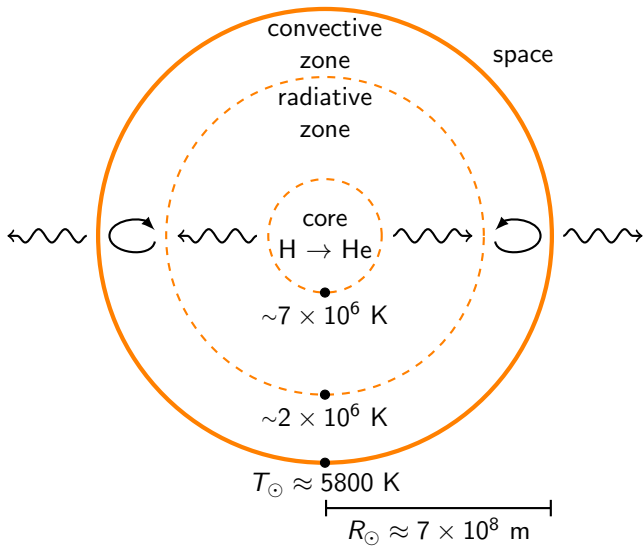
They Might Be Giants (1993): [Why Does the Sun Shine?](#)

# The sun is a miasma of incandescent **plasma**

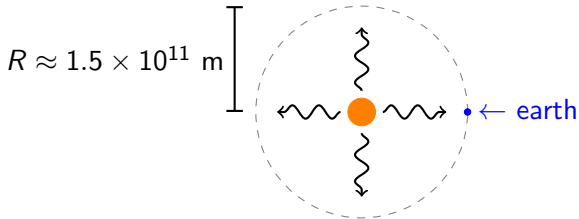


They Might Be Giants (2009): [Why Does the Sun Really Shine?](#)

# The sun



# Sunlight in space



- as a black body at  $T_{\odot} \approx 5800 \text{ K}$ , sun's surface radiates

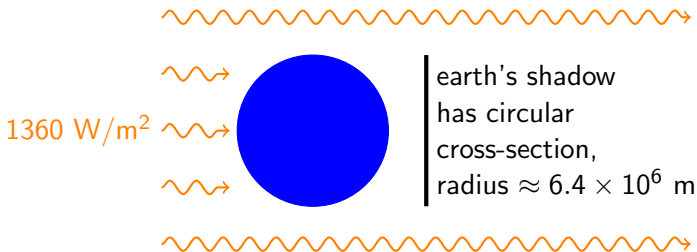
$$\sigma A_{\odot} T_{\odot}^4 = 4\sigma\pi R_{\odot}^2 T_{\odot}^4 \approx 4 \times 10^{26} \text{ W}$$

- at earth's distance from sun, irradiance is

$$S_0 := \frac{\sigma A_{\odot} T_{\odot}^4}{4\pi R^2} \approx \frac{4 \times 10^{26} \text{ W}}{4\pi(1.5 \times 10^{11} \text{ m})^2} \approx 1360 \text{ W/m}^2$$

- **solar constant**  $S_0$  is irradiance at top of earth's atmosphere

# How much solar power 'hits' the earth?



- solar power incident on earth's upper atmosphere:

$$(1360 \text{ W/m}^2)\pi(6.4 \times 10^6 \text{ m})^2 \approx 1.7 \times 10^{17} \text{ W}$$

- $\sim 10,000$  times humanity's  $\sim 1.8 \times 10^{13}$  W used for all purposes

# (Almost) all energy on earth is (?) solar energy

- hydro  $\leftarrow$  solar (via the water cycle)
- wind  $\leftarrow$  solar (via hot air rising)
- biomass  $\leftarrow$  solar (via photosynthesis)
- fossil fuels  $\leftarrow$  biomass (via rotting/cooking underground)  $\leftarrow$  solar

but

- nuclear  $\not\leftarrow$  solar
- geothermal  $\leftarrow$  nuclear (via reactions underground)

and solar  $\leftarrow$  nuclear

so really, all (?) energy on earth is **nuclear** energy

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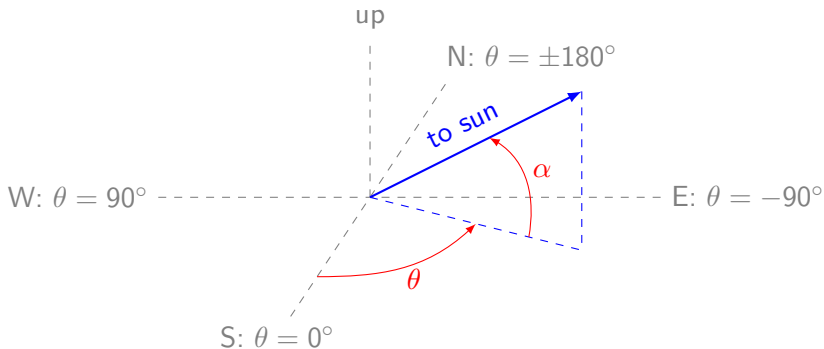
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# Elevation and azimuth angles



- origin is at observer on earth's surface
- observer sees sun at **elevation**  $\alpha$  and **azimuth**  $\theta$
- $\alpha = 0^\circ$  at sunrise/sunset
- $\theta = 0^\circ$  if sun is due south, increases clockwise

# Elevation and azimuth formulas

$\alpha =$

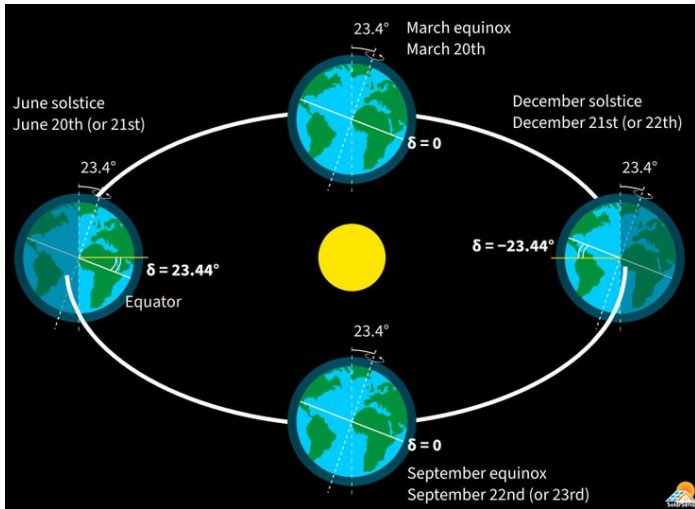
$\theta =$

- $\phi_0$  is observer's **latitude**
- $\delta$  and  $\lambda$  are **declination** and **hour** angles
- $\text{acos}$  returns values in  $[0^\circ, 90^\circ]$ ;  $\text{atan2}$  in  $[-180^\circ, 180^\circ]$

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Zhang (2021): A solar azimuth formula based on the subsolar point and  $\text{atan2}$

# Declination angle



$\delta \approx$

,  $\gamma =$

, day #  $d$  is

# Hour angle

- the empirical **equation of time**,

$$\tau \approx$$

converts Greenwich Mean Time  $t_{\text{gm}}$  to local **solar time** (solar time,  $t_{\text{gm}} + \tau$ , equals 12 h when sun is highest)

- the hour angle,

$$\lambda =$$

is defined such that

- ◇  $\lambda$  equals observer's **longitude**  $\lambda_0$  when sun is highest
- ◇  $\lambda$  increases by  $360^\circ$  every 24 h

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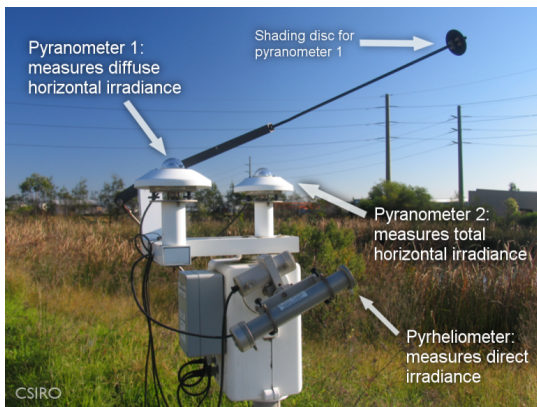
**Incident irradiance on surfaces**

Solar photovoltaics

# Sunlight at earth's surface

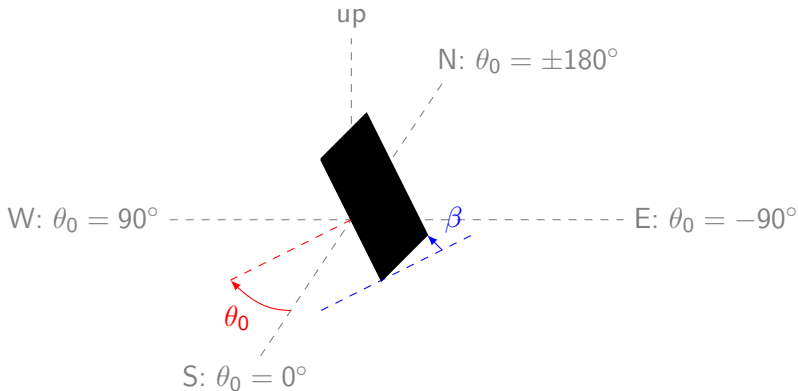
- on a clear day,  $\sim 75\%$  of  $S_0$  reaches earth's surface
  - ◇  $\sim 70\%$  transmitted through atmosphere (**beam**)
  - ◇  $\sim 5\%$  scattered to earth by air, dust, water vapor (**diffuse**)
  - ◇  $\sim 20\%$  absorbed by atmosphere
  - ◇  $\sim 5\%$  scattered back to space
- as cloud cover increases,
  - ◇ less of  $S_0$  reaches surface (as little as  $\sim 10\%$ )
  - ◇ beam % of total surface sunlight falls, diffuse % rises

# Sunlight measurements



- weather data services often provide one or more of
  - ◇ beam irradiance  $S_b^\perp$  on surface  $\perp$  to sunbeam
  - ◇ diffuse irradiance  $S_d^-$  on horizontal surface
  - ◇ total irradiance  $S_{\text{tot}}^-$  on horizontal surface
  - ◇ beam irradiance  $S_b^-$  on horizontal surface

# Surface tilt and azimuth



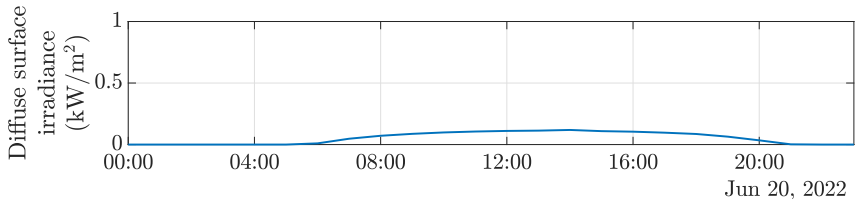
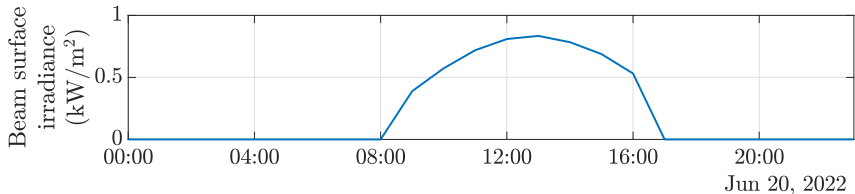
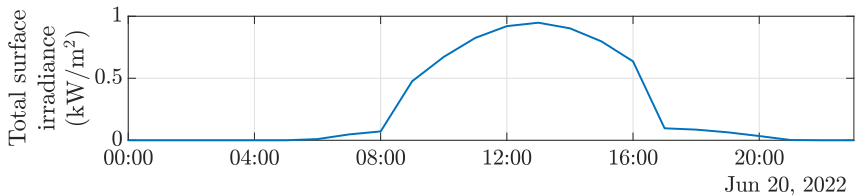
- surface **tilt**  $\beta$  and **azimuth**  $\theta_0$  define its orientation
- $\beta = 0^\circ$  for horizontal surfaces,  $90^\circ$  for vertical
- $\theta_0$  follows same convention as sun's azimuth  $\theta$

# Irradiance on arbitrarily oriented surface

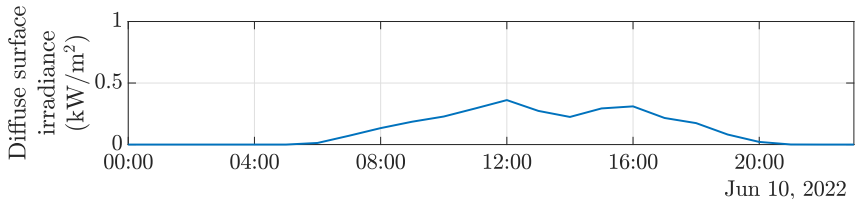
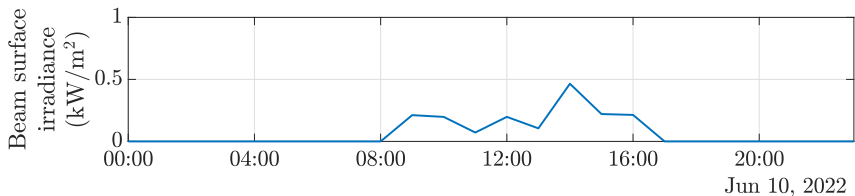
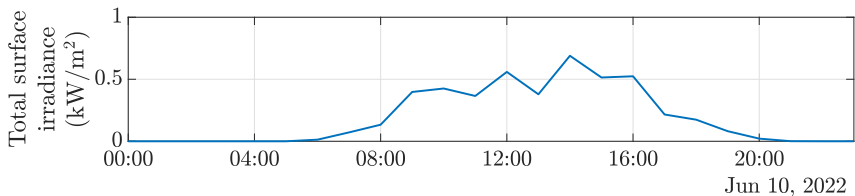
$$S_b = \begin{cases} S_d \approx \\ S_{\text{tot}} = \end{cases}$$

- $S_b$  and  $S_{\text{tot}}$  are beam and total irradiance on surface
- $\alpha < 0^\circ$  means sun is down
- $90^\circ < |\theta - \theta_0| < 270^\circ$  means sun is behind surface

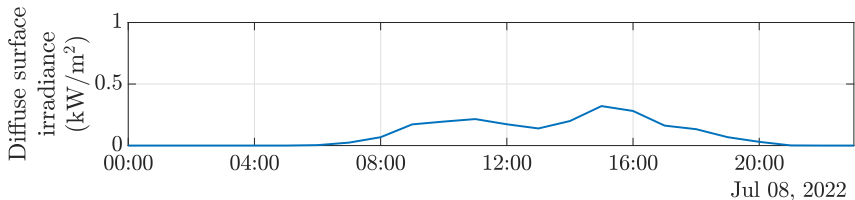
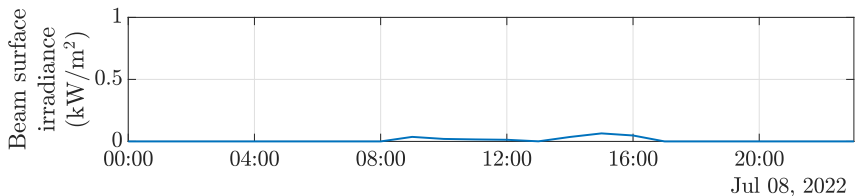
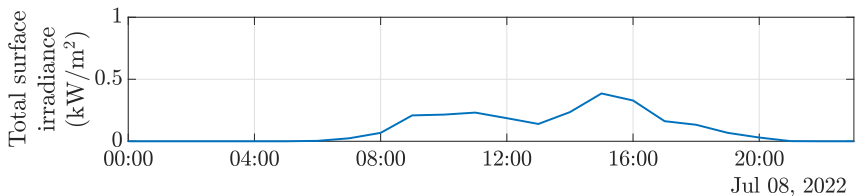
# Clear summer day with $\beta = \phi_0$ , $\theta_0 = 0^\circ$



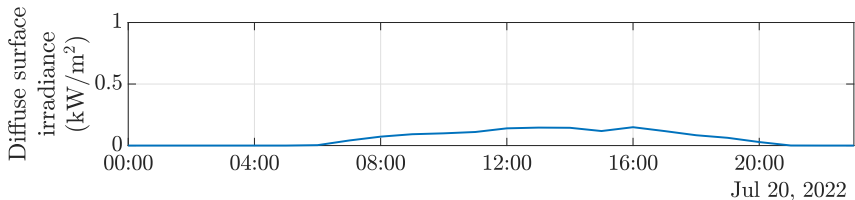
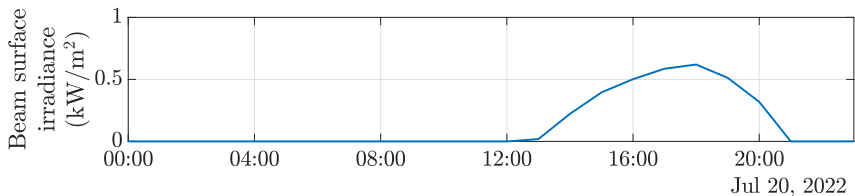
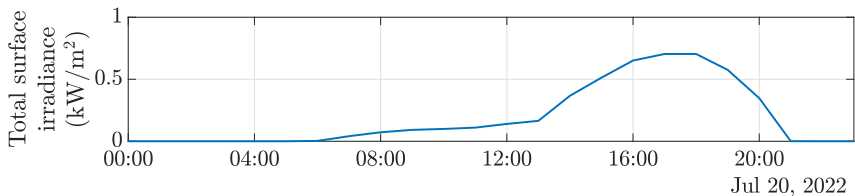
# Partly cloudy summer day with $\beta = \phi_0$ , $\theta_0 = 0^\circ$



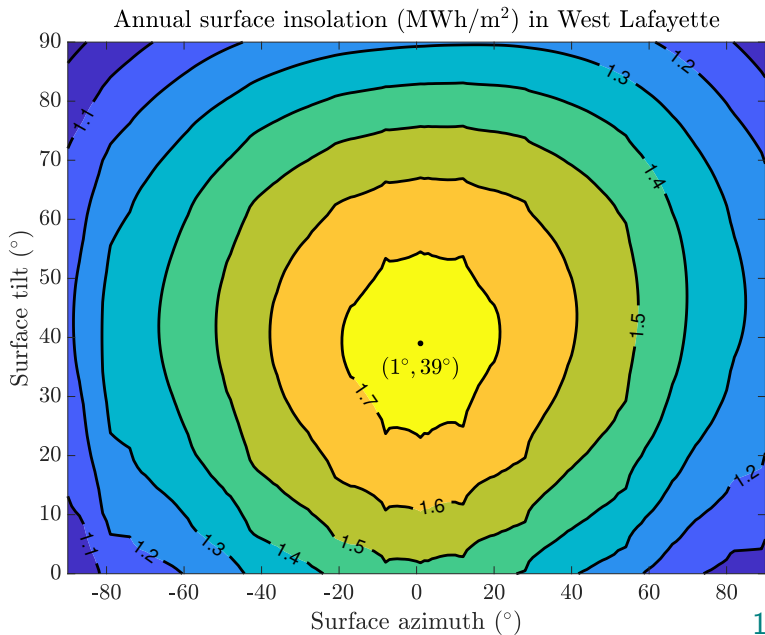
# Cloudy summer day with $\beta = \phi_0$ , $\theta_0 = 0^\circ$



# Clear summer day with $\beta = \theta_0 = 90^\circ$



$\beta \approx \phi_0$  and  $\theta_0 \approx 0^\circ$  maximize annual incident energy



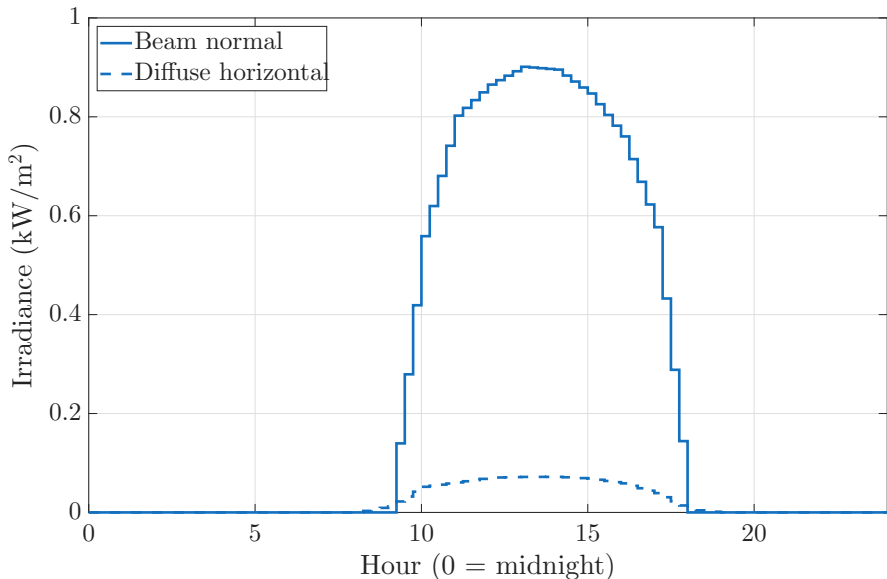
# Sunlight and windows

- most windows are vertical:  $\beta = 90^\circ$
- with  $\sin(\beta) = 1$  and  $\cos(\beta) = 0$ , incident irradiance reduces to

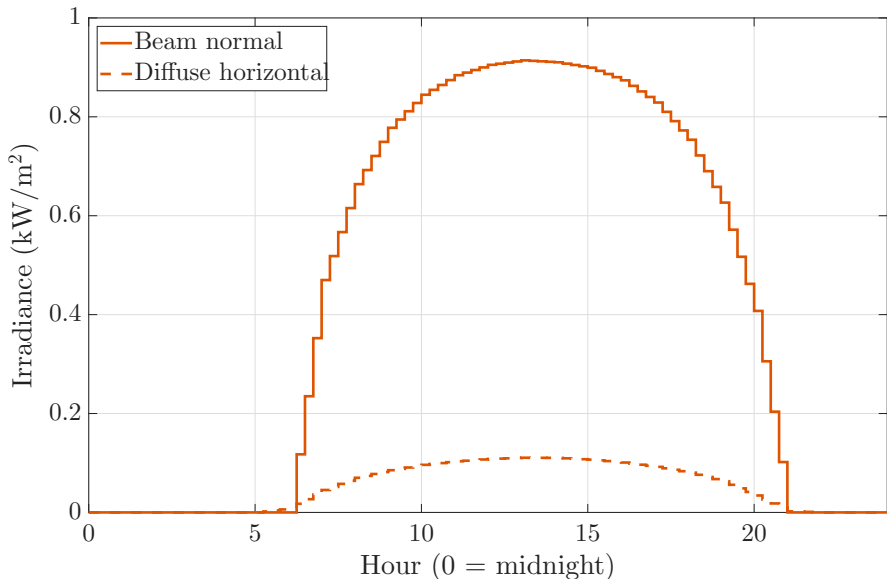
$$S_{\text{tot}} \approx \left\{ \right.$$

- a window of area  $A$  transmits solar power  $cAS_{\text{tot}}$
- $c \in [0, 1]$  is the window's **solar heat gain coefficient** (typically,  $c \approx 0.25$  to  $0.8$ )
- can simulate shading (from trees, blinds, ...) by adjusting  $c$

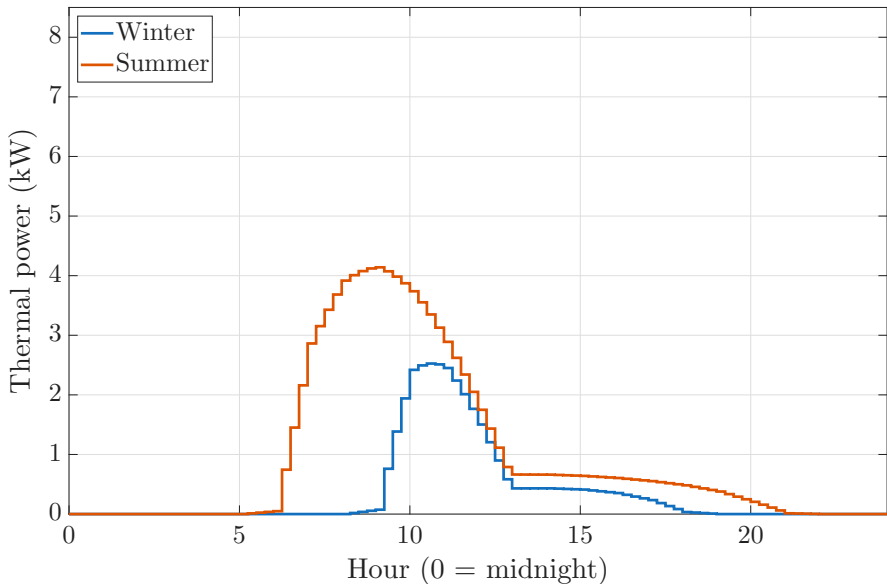
## Winter irradiance data for window examples



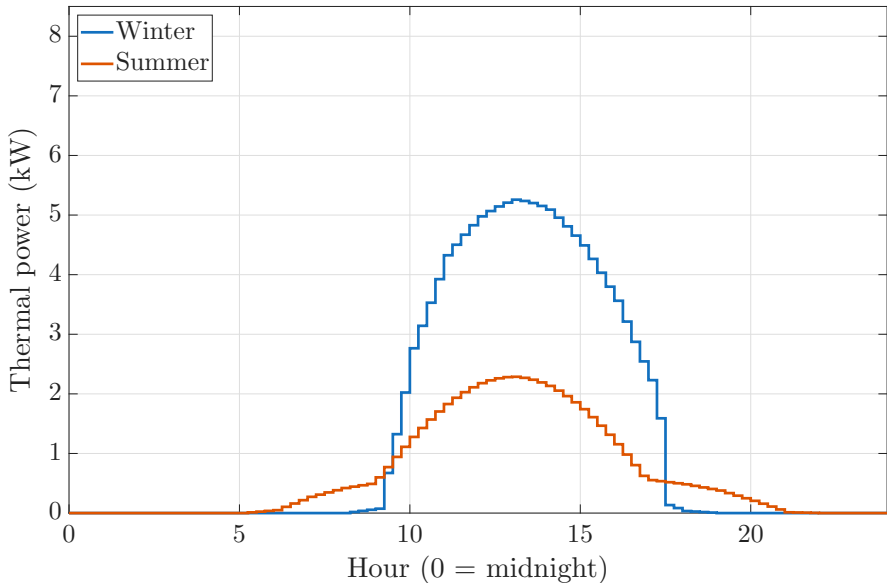
## Summer irradiance data for window examples



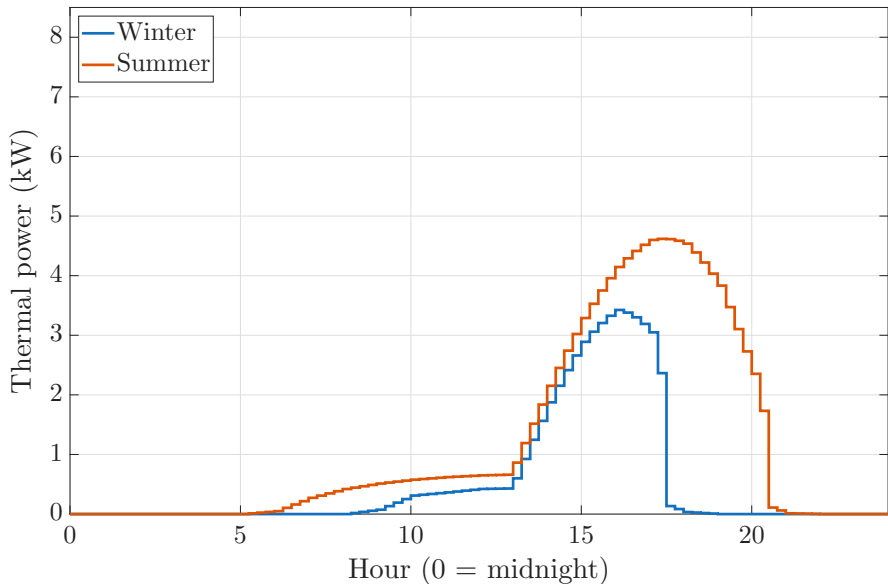
# Solar thermal power through east-facing windows



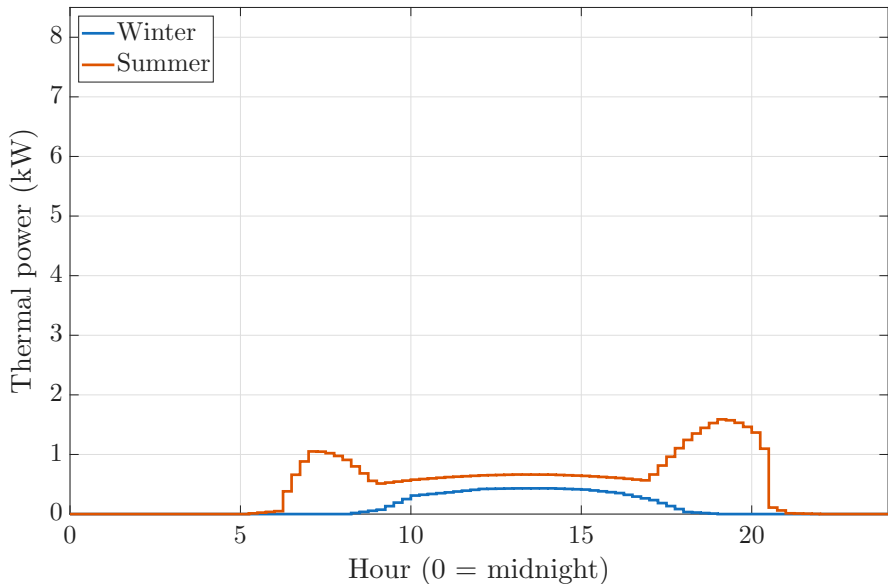
# Solar thermal power through **south**-facing windows



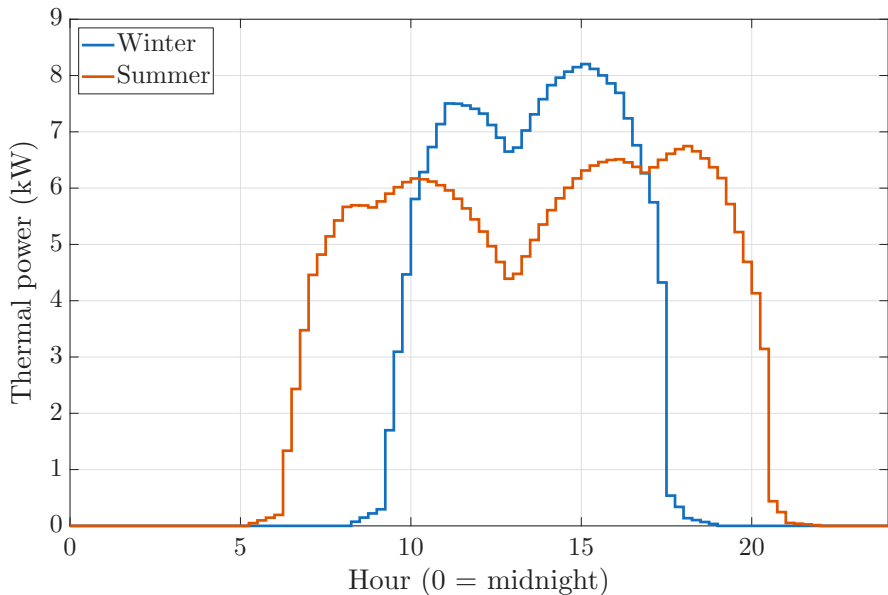
# Solar thermal power through west-facing windows



# Solar thermal power through **north**-facing windows



# Solar thermal power through **all** windows



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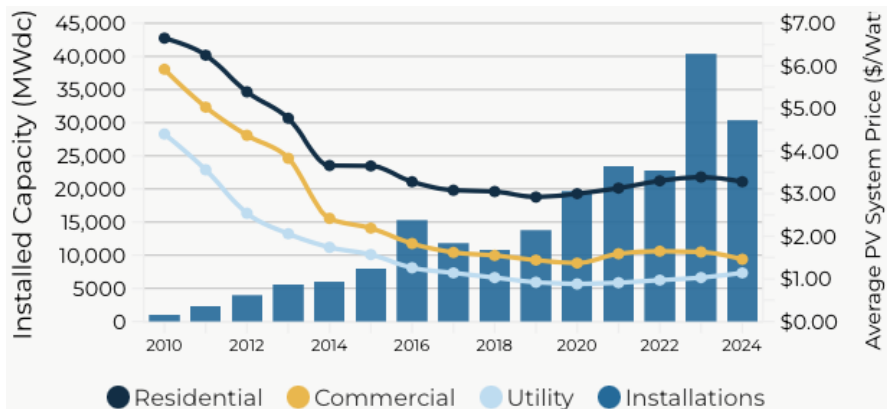
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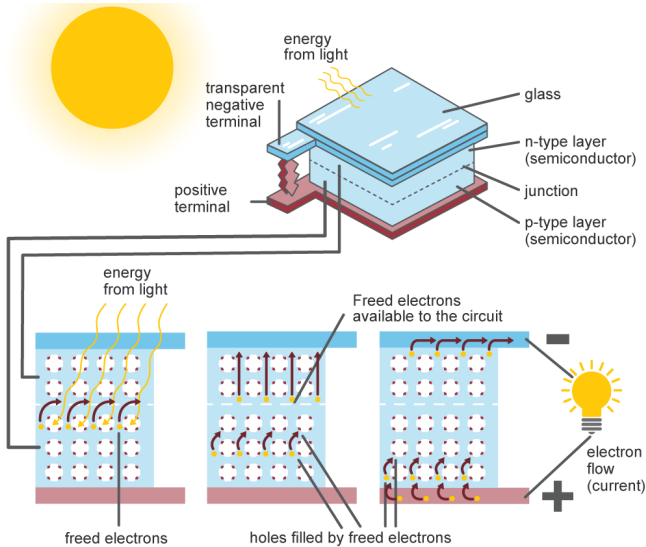
**Solar photovoltaics**

# US costs and installations of solar photovoltaics



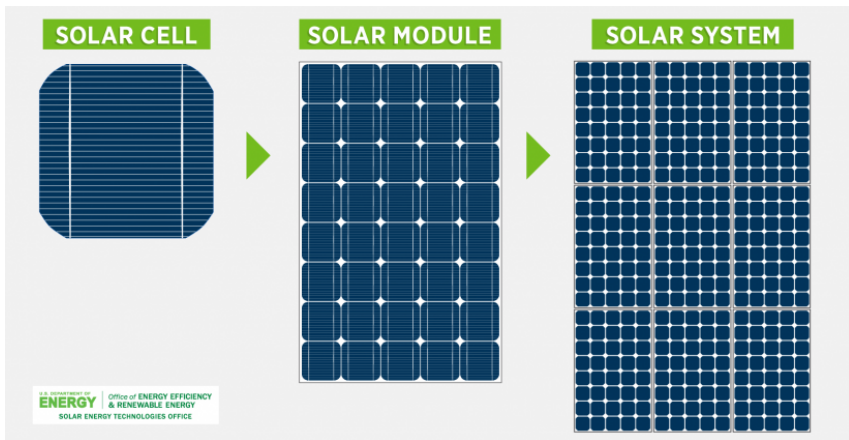
Solar Energy Industries Association (2024): [Solar Industry Research Data](#)

# Solar photovoltaic cells



US Energy Information Administration: [Photovoltaics and electricity](#)

# Solar photovoltaic modules and arrays



- typical cell size:  $\sim 6'' \times 6''$ ,  $\sim 0.023 \text{ m}^2$ ,  $\sim 4.2 \text{ W peak}$
- typical module size:  $\sim 5' \times 3'$ ,  $\sim 1.4 \text{ m}^2$ ,  $\sim 250 \text{ W peak}$

# Solar photovoltaic efficiency

$$\eta = \frac{\text{electric power output}}{\text{radiative power input}}$$

- whole-system efficiency includes cells, conductors, inverter, ...
- for typical solar arrays,  $\eta \approx 18$  to  $22\%$
- Shockley-Queisser limit:  $\eta \leq \sim 33\%$  for any single-junction cell

# Efficiency and temperature

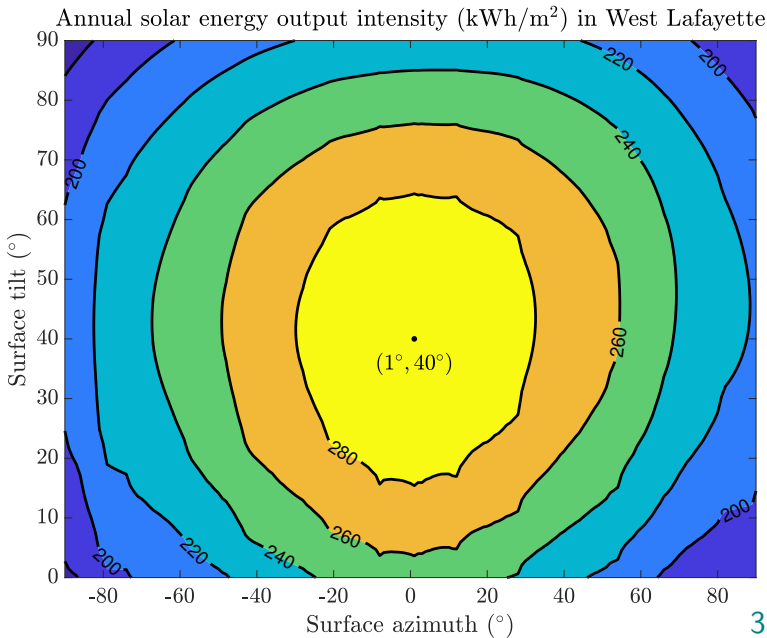
- efficiency scales  $\sim$ linearly with cell temperature  $T_c$

$$\eta \approx$$

- $\tilde{\eta}$  is efficiency at rated cell temperature  $\tilde{T} \approx 25 \text{ }^\circ\text{C}$
- $T_0 \approx 270 \text{ }^\circ\text{C}$  is cell temperature at which generation stops
- cell temperature scales  $\sim$ linearly with outdoor temperature  $T_a$

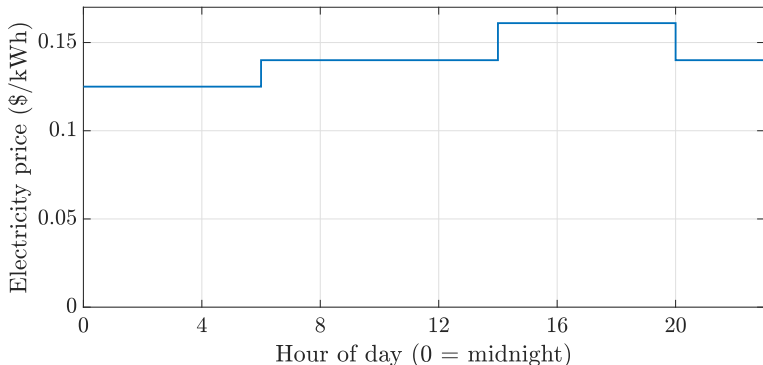
$$T_c \approx$$

$\beta \approx \phi_0$  and  $\theta_0 \approx 0^\circ$  maximize annual energy output

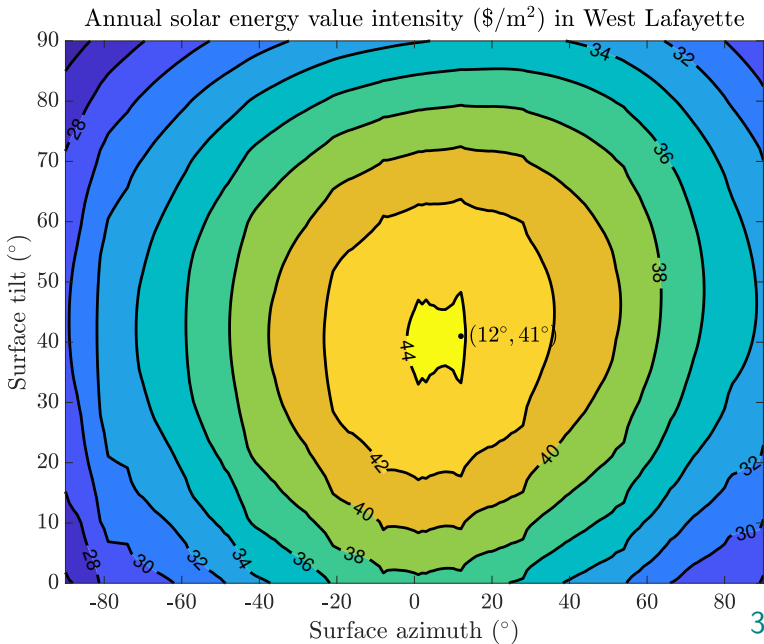


# Time-varying electricity prices

- most people buy electricity at a constant price
- some see 'time-of-use' prices with 2 or 3 tiers
- a few see hourly prices tied to wholesale markets



# Time-of-use pricing shifts optimal panel orientation west



# Net metering

- sometimes, rooftop solar supply exceeds building demand
- some utilities buy excess power at their electricity sale price
- others pay a lower price; some pay nothing at all

